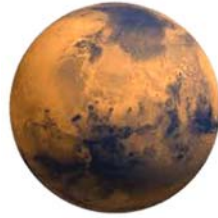


Mars' Rotation Equations

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Constants

$f_R := 2791.826$	Free Rotation Constant
$i_M := 0.127036$	Maximum Influenced Rotation Constant (for planets and moons only)
$i_{St} := 1.0121647 \cdot 10^{-12}$	Start Influenced Rotation Distance Constant
$i_{Ma} := 5.6964797 \cdot 10^{-10}$	Maximum Influenced Rotation Distance Constant
$i_{Sp} := 1.0686849 \cdot 10^{-9}$	Stop Rotation Distance Constant

Facts

	<u>Mars</u>	<u>Sun</u>
Mass (kg)	$\underline{m} := 6.4185 \cdot 10^{23}$	$M := 1.9891 \cdot 10^{30}$
Density (g/cm ³)	$\rho := 3.94$	$\rho_s := 1.408$
Axis Tilt (deg)	$t := 25.19$	$t_s := 7.25$
Semi-major Axis (km)	$a := 227920000$	
Orbit Eccentricity (deg)	$\underline{e} := 0.09341233$	
Orbit Inclination (degree), with respect to equator	$i := 5.66$	

$$\omega_F := f_R \div \sqrt[6]{m} \cdot \sqrt[2]{\rho}$$

$$\omega_F = 0.59666542$$

Mars' Free Rotation (per day)

Part 1

Mars' Influenced Rotation by the influence of the Sun



$$q := a \cdot (1 - e)$$

$$q = 206629461.7 \quad \text{Mars' Perihelion Distance (km)}$$

$$Q := a \cdot (1 + e)$$

$$Q = 249210538.3 \quad \text{Mars' Aphelion Distance (km)}$$

$$i_r := \left(\left| \cos\left(\frac{i \cdot \pi}{180}\right) \right| + 1 \right) \div 2$$

$$i_r = 0.99756233 \quad \text{Mars' Influenced Rotation Reduction Factor by Orbit Inclination}$$

$$\omega_{Mi} := \frac{\sqrt[6]{m \cdot i_r \div M} \div \sqrt[6]{\rho}}{i_M}$$

$$\omega_{Mi} = 0.51853329 \quad \text{Mars' Maximum Influenced Rotation by the Sun (p.d.)}$$

$$S_t := \frac{\sqrt[6]{m \cdot i_r \div M}}{i_{St}}$$

$$S_t = 81790269932.4 \quad \text{Mars' Start Influenced Rotation Distance to the Sun (km)}$$

$$M_a := \frac{\sqrt[6]{m \cdot i_r \div M}}{i_{Ma}}$$

$$M_a = 145326988.6 \quad \text{Mars' Maximum Influenced Rotation Distance to the Sun (km)}$$

$$S_p := \frac{\sqrt[6]{m \cdot i_r \div M}}{i_{Sp}}$$

$$S_p = 77464577.3 \quad \text{Mars' Stop Rotation Distance to the Sun (km)}$$

Calculating Mars' average distance to the Sun, if ($q < S_p < Q$)

$$x := \text{if} \left(q < S_p, \text{if} \left(S_p < Q, \frac{S_p - a}{e}, 0 \right), 0 \right)$$

$x = 0$ X value at Mars' orbit intersection with S_p Boundary (km)

$$b := a\sqrt{1 - e^2}$$

$b = 226923421.8$ Mars' Semi-minor Axis (km)

$$y := b\sqrt{a^2 - x^2} \div a$$

$y = 226923421.84$ Y value at the Mars' orbit intersection with S_p Boundary (km)

$$\theta := \text{atan} \left(\frac{-x}{y} \right) + \frac{\pi}{2}$$

$\theta = 1.57079633$ Half-angle of the Mars' orbit out of S_p Boundary (rad)

$$s := a \cdot \int_0^\theta \sqrt{1 - e^2 \cdot \sin(\theta)^2} d\theta$$

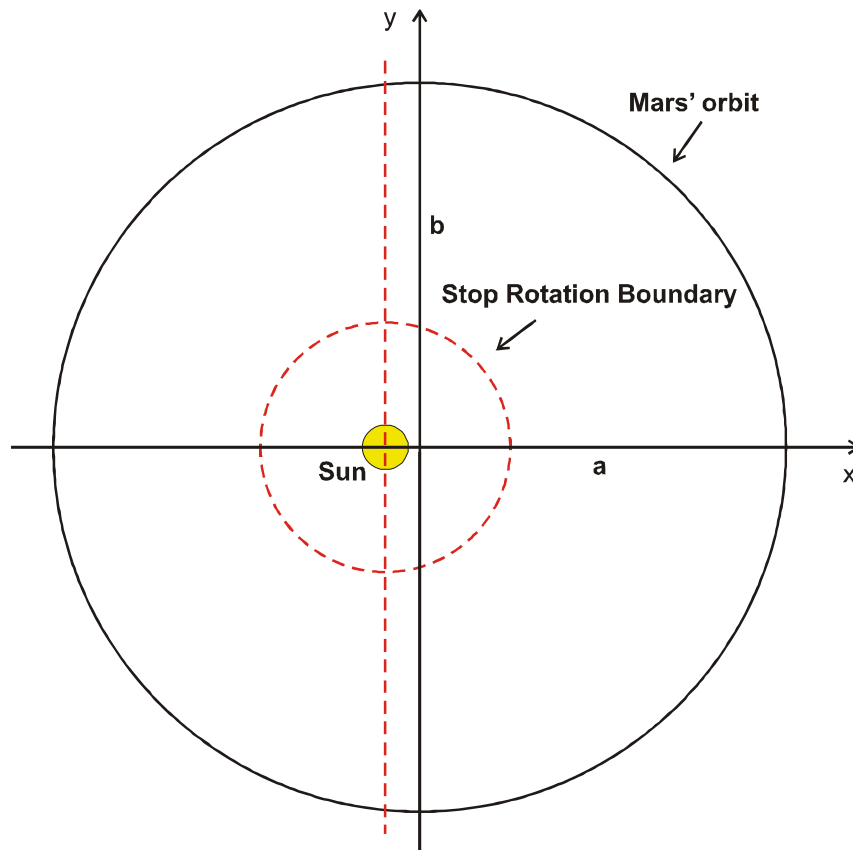
$s = 357233616.89$ Half of Mars' orbit out of S_p Boundary (km)

$$a_a := \text{if} \left[q < S_p, \text{if} \left[S_p < Q, a \frac{\int_{\pi - \frac{s}{a}}^\pi (1 - e \cdot \cos(E)) \cdot \sqrt{1 - e^2 \cdot \cos(E)^2} dE}{\int_{\pi - \frac{s}{a}}^\pi \sqrt{1 - e^2 \cdot \cos(E)^2} dE}, 0 \right], 0 \right]$$

$a_a = 0$ Mars' average distance to the Sun outside S_p Boundary (km)

$$n := \frac{2 \cdot s}{2 \cdot a \cdot \int_0^\pi \sqrt{1 - e^2 \cdot \sin(\theta)^2} d\theta} \cdot \sqrt{\frac{a_a^3}{a^3}}$$

$n = 0$ Ratio of the Mars' orbit out of S_p Boundary to the whole orbit



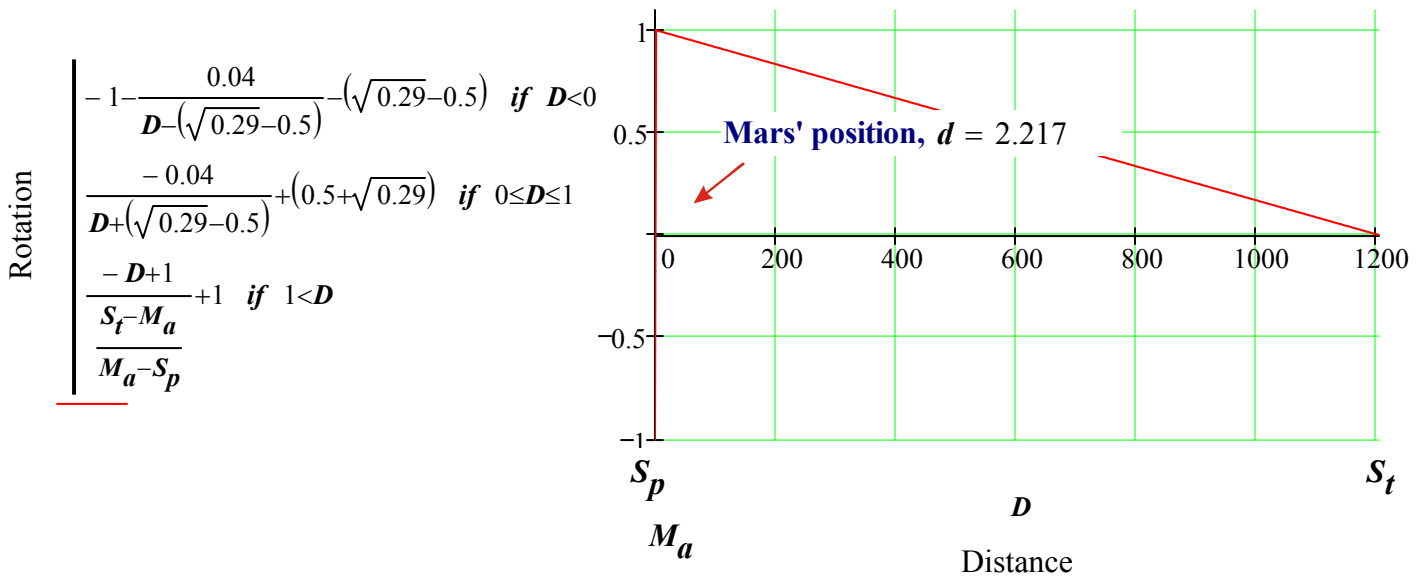
Mars' orbit relative to the Stop Rotation Boundary

$$d := \text{if} \left(q < S_p, \text{if} \left(S_p < Q, \frac{a_a - S_p}{M_a - S_p}, \frac{a - S_p}{M_a - S_p} \right), \frac{a - S_p}{M_a - S_p} \right)$$

$$d = 2.21706568$$

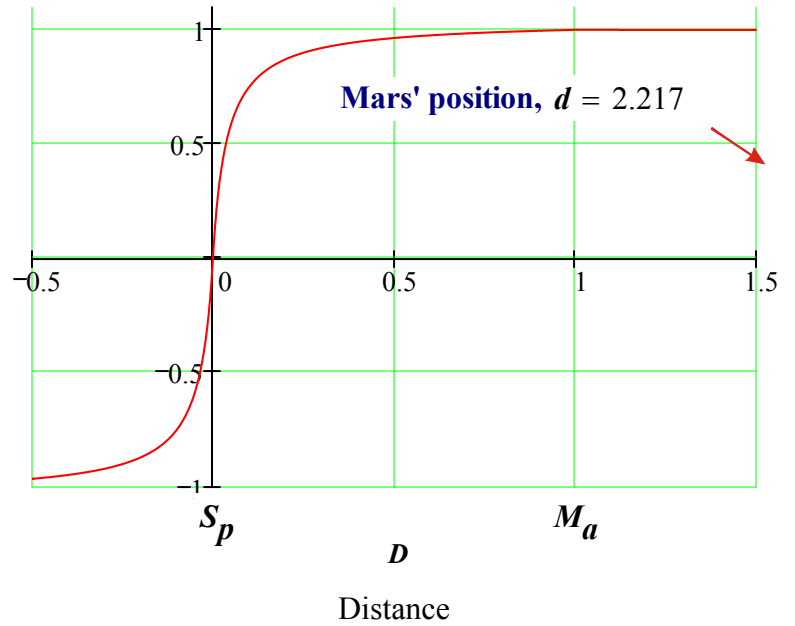
Mars' corresponding distance to the Sun relative to S_p on the X axis of the graph

Non-proportional Rotation Graph



$$\text{Rotation} = \begin{cases} -1 - \frac{0.04}{D - (\sqrt{0.29} - 0.5)} - (\sqrt{0.29} - 0.5) & \text{if } D < 0 \\ \frac{-0.04}{D + (\sqrt{0.29} - 0.5)} + (0.5 + \sqrt{0.29}) & \text{if } 0 \leq D \leq 1 \\ \frac{-D+1}{S_t - M_a} + 1 & \text{if } 1 < D \\ \frac{M_a - S_p}{M_a - S_p} & \end{cases}$$

Left end of the Rotation Graph



$$\omega(d) := \begin{cases} -1 - \frac{0.04}{d - (\sqrt{0.29} - 0.5)} - (\sqrt{0.29} - 0.5) & \text{if } d < 0 \\ \frac{-0.04}{d + (\sqrt{0.29} - 0.5)} + (0.5 + \sqrt{0.29}) & \text{if } 0 \leq d \leq 1 \\ \frac{-d+1}{S_t - M_a} + 1 & \text{if } 1 < d \\ \frac{M_a - S_p}{M_a - S_p} & \end{cases}$$

$\omega(d) = 0.99898839$ Mars' corresponding Influenced Rotation by the Sun on the Y axis of the graph

$$t_r := \text{if} \left(a < M_a, \text{if} \left(\omega_{Mi} > \omega_F, \frac{t \cdot \omega_F}{90}, \frac{t \cdot \omega_{Mi}}{90} \right), \text{if} \left(\omega(d) \cdot \omega_{Mi} > \omega_F, \frac{t \cdot \omega_F}{90}, \frac{t \cdot \omega(d) \cdot \omega_{Mi}}{90} \right) \right)$$

$t_r = 0.14498489$ Mars' Maximum and Free Rotational Speed Reduction by Axis Tilt

$$\omega_1 := \text{if} \left[a > M_a, \omega(d) \cdot \omega_{Mi} + \omega_F - t_r, \left[\omega(d) \cdot (\omega_{Mi} + \omega_F - t_r) \cdot \text{if} \left(q < S_p, \text{if} \left(Q > S_p, n, 0 \right), 1 \right) \right] \right]$$

$\omega_1 = 0.96968927$ Mars' end result Rotation (p.d.)

Part 2

Mars' Influenced Rotation by the influence of the Phobos

if ($q < S_t$)



Phobos' Facts

$a_m := 9378$	Phobos Semi-major Axis (km)
$e_m := 0.0151$	Phobos Orbit Eccentricity (degree)
$i_m := 1.08$	Phobos Orbit Inclination (degree)
$t_m := 0.046$	Phobos Axis Tilt (degree)
$m_m := 1.06 \cdot 10^{16}$	Phobos Mass (kg)

$$q := a_m \cdot (1 - e_m)$$

$q = 9236.4$ Phobos' Perihelion Distance (km)

$$Q := a_m \cdot (1 + e_m)$$

$Q = 9519.6$ Phobos' Aphelion Distance (km)

$$i_r := \left(\left| \cos \left(\frac{t_m \cdot \pi}{180} \right) \right| + 1 \right) \div 2$$

$i_r = 0.99999984$ Phobos' Orbit Inclination Reduction Factor

$$\omega_{Mi} := \sqrt[6]{m_m \cdot i_r \div m} \div \sqrt[6]{\rho} \div i_M \div \sqrt{M \div m_m}$$

$\omega_{Mi} = 0.000000023$ Mars' Maximum Influenced Rotation by the Phobos (p.d.)

$$S_t := \sqrt[6]{m_m \cdot i_r \div m} \div i_{St} \div \sqrt{M \div m_m}$$

$S_t = 3639.6$ Mars' Start Influenced Rotation Distance to the Phobos (km)

$$M_a := \sqrt[6]{m_m \cdot i_r \div m} \div i_{Ma} \div \sqrt{M \div m_m}$$

$M_a = 6.5$ Mars' Maximum Influenced Rotation Distance to the Phobos (km)

$$S_p := \sqrt[6]{m_m \cdot i_r \div m} \div i_{Sp} \div \sqrt{M \div m_m}$$

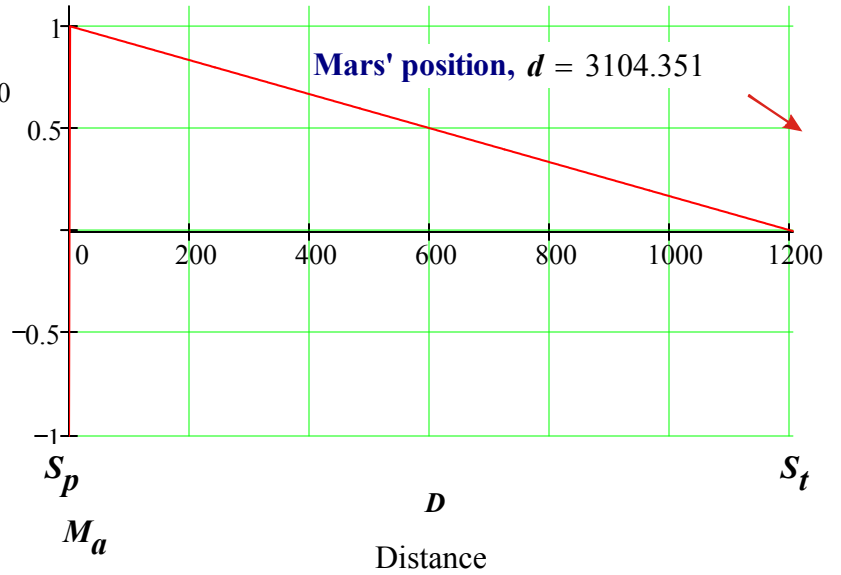
$S_p = 3.4$ Mars' Stop Rotation Distance to the Phobos (km)

$$d := \frac{a_m - S_p}{M_a - S_p}$$

$d = 3.10435067 \times 10^3$ Mars' corresponding distance to the Phobos relative to S_p on the X axis of the graph

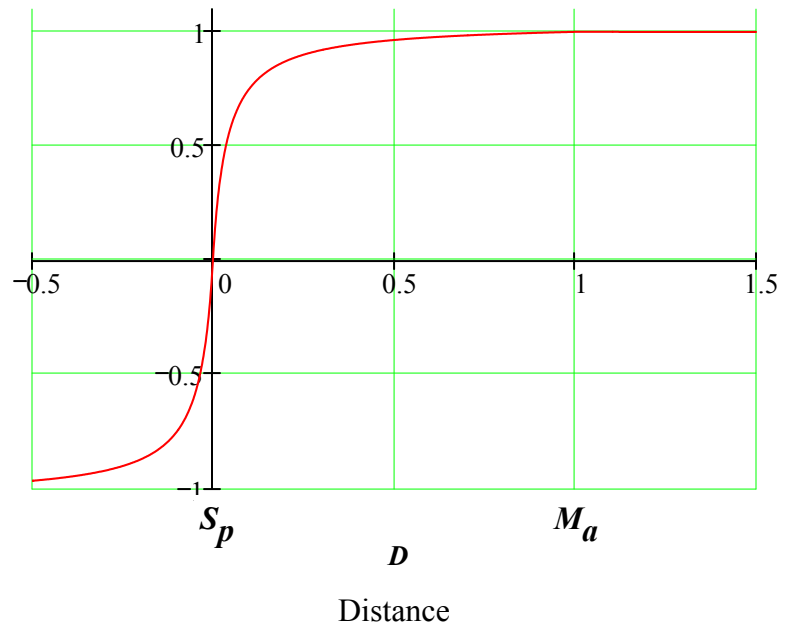
$$\text{Rotation} = \begin{cases} -1 - \frac{0.04}{D - (\sqrt{0.29} - 0.5)} - (\sqrt{0.29} - 0.5) & \text{if } D < 0 \\ \frac{-0.04}{D + (\sqrt{0.29} - 0.5)} + (0.5 + \sqrt{0.29}) & \text{if } 0 \leq D \leq 1 \\ \frac{-D+1}{S_t - M_a} + 1 & \text{if } 1 < D \\ \frac{M_a - S_p}{M_a - S_p} & \end{cases}$$

Non-proportional Rotation Graph



$$\text{Rotation} = \begin{cases} -1 - \frac{0.04}{D - (\sqrt{0.29} - 0.5)} - (\sqrt{0.29} - 0.5) & \text{if } D < 0 \\ \frac{-0.04}{D + (\sqrt{0.29} - 0.5)} + (0.5 + \sqrt{0.29}) & \text{if } 0 \leq D \leq 1 \\ \frac{-D+1}{S_t - M_a} + 1 & \text{if } 1 < D \\ \frac{M_a - S_p}{M_a - S_p} & \end{cases}$$

Left end of the Rotation Graph



$$\omega(d) := \begin{cases} -1 \cdot \left(\frac{S_p}{M_a - S_p} \right) - \frac{0.04 \cdot \left(\frac{S_p}{M_a - S_p} \right)}{d - (\sqrt{0.29} - 0.5)} - (\sqrt{0.29} - 0.5) & \text{if } d < 0 \\ \frac{-0.04}{d + (\sqrt{0.29} - 0.5)} + (0.5 + \sqrt{0.29}) & \text{if } 0 \leq d \leq 1 \\ \frac{-d + 1}{\frac{S_t - M_a}{M_a - S_p}} + 1 & \text{if } 1 < d \end{cases}$$

$\omega(d) = -1.57947219$ Mars' corresponding Influenced Rotation by the Phobos on the Y axis of the graph

$$t_r := \text{if} \left(a_m < M_a, \text{if} \left(\omega_{Mi} > \omega_F, \frac{i_m \cdot \omega_F}{90}, \frac{i_m \cdot \omega_{Mi}}{90} \right), \text{if} \left(\omega(d) \cdot \omega_{Mi} > \omega_F, \frac{i_m \cdot \omega_F}{90}, \frac{i_m \cdot \omega(d) \cdot \omega_{Mi}}{90} \right) \right)$$

$t_r = -4.3734119 \times 10^{-10}$ Mars' Maximum and Free Rotational Speed Reduction by Axis Tilt

$$\omega i_2 := \text{if} \left[a_m < M_a, \omega(d) \cdot (\omega_{Mi} - t_r), \text{if} \left(q < S_t, \omega(d) \cdot \omega_{Mi} - t_r, 0 \right) \right]$$

$\omega i_2 = 0$

Mars' end result Influenced Rotation by the Phobos (p.d.)

(Negative number means the reduction amount from Mars' Free Rotation)

Part 3

Mars' Influenced Rotation by the influence of the Deimos



Deimos' Facts

a_{mm}	:= 23459	Deimos Semi-major Axis (km)
e_{mm}	:= 0.0002	Deimos Orbit Eccentricity (degree)
i_{mm}	:= 1.79	Deimos Orbit Inclination (degree)
t_{mm}	:= 0.897	Deimos Axis Tilt (degree)
m_{mm}	:= $2.4 \cdot 10^{15}$	Deimos Mass (kg)

$$q := a_m \cdot (1 - e_m)$$

$q = 23454.3$ Deimos' Perihelion Distance (km)

$$Q := a_m \cdot (1 + e_m)$$

$Q = 23463.7$ Deimos' Aphelion Distance (km)

$$i_r := \left(\left| \cos \left(\frac{t_m \cdot \pi}{180} \right) \right| + 1 \right) \div 2$$

$i_r = 0.99993873$ Deimos' Orbit Inclination Reduction Factor

$$\omega_{Mi} := \sqrt[6]{m_m \cdot i_r \div m} \div \sqrt[6]{\rho} \div i_M \div \sqrt{M \div m_m}$$

$\omega_{Mi} = 0.000000009$ Mars' Maximum Influenced Rotation by the Deimos (p.d.)

$$S_t := \sqrt[6]{m_m \cdot i_r \div m} \div i_{St} \div \sqrt{M \div m_m}$$

$S_t = 1352$ Mars' Start Influenced Rotation Distance to the Deimos (km)

$$M_a := \sqrt[6]{m_m \cdot i_r \div m} \div i_{Ma} \div \sqrt{M \div m_m}$$

$M_a = 2.4$ Mars' Maximum Influenced Rotation Distance to the Deimos (km)

$$S_p := \sqrt[6]{m_m \cdot i_r \div m} \div i_{Sp} \div \sqrt{M \div m_m}$$

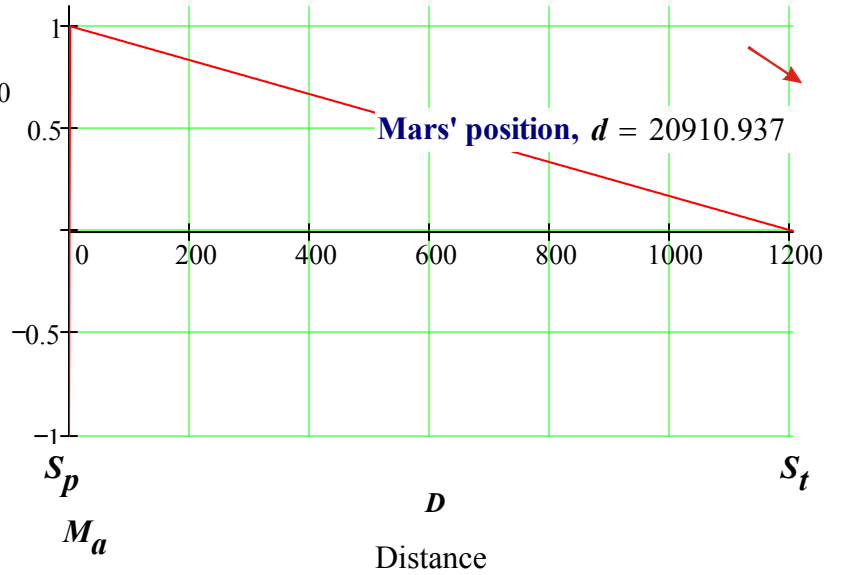
$S_p = 1.3$ Mars' Stop Rotation Distance to the Deimos (km)

$$d_{\text{ww}} := \frac{a_m - S_p}{M_a - S_p}$$

$d = 2.09109375 \times 10^4$ Mars' corresponding distance to the Deimos relative to S_p on the X axis of the graph

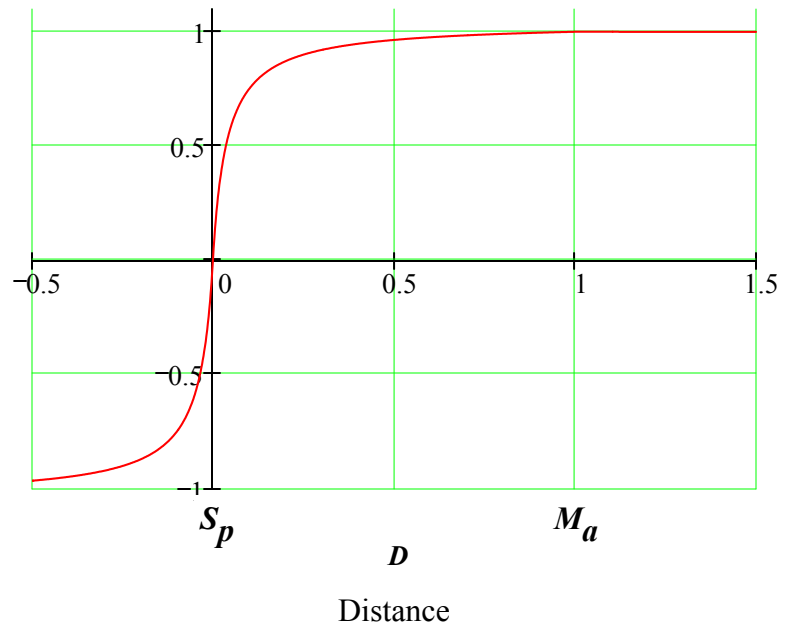
$$\text{Rotation} = \begin{cases} -1 - \frac{0.04}{D - (\sqrt{0.29} - 0.5)} - (\sqrt{0.29} - 0.5) & \text{if } D < 0 \\ \frac{-0.04}{D + (\sqrt{0.29} - 0.5)} + (0.5 + \sqrt{0.29}) & \text{if } 0 \leq D \leq 1 \\ \frac{-D+1}{S_t - M_a} + 1 & \text{if } 1 < D \\ \frac{M_a - S_p}{M_a - S_p} & \end{cases}$$

Non-proportional Rotation Graph



$$\text{Rotation} = \begin{cases} -1 - \frac{0.04}{D - (\sqrt{0.29} - 0.5)} - (\sqrt{0.29} - 0.5) & \text{if } D < 0 \\ \frac{-0.04}{D + (\sqrt{0.29} - 0.5)} + (0.5 + \sqrt{0.29}) & \text{if } 0 \leq D \leq 1 \\ \frac{-D+1}{S_t - M_a} + 1 & \text{if } 1 < D \\ \frac{M_a - S_p}{M_a - S_p} & \end{cases}$$

Left end of the Rotation Graph



$$\omega(d) := \begin{cases} -1 \cdot \left(\frac{S_p}{M_a - S_p} \right) - \frac{0.04 \cdot \left(\frac{S_p}{M_a - S_p} \right)}{d - (\sqrt{0.29} - 0.5)} - (\sqrt{0.29} - 0.5) & \text{if } d < 0 \\ \frac{-0.04}{d + (\sqrt{0.29} - 0.5)} + (0.5 + \sqrt{0.29}) & \text{if } 0 \leq d \leq 1 \\ \frac{-d + 1}{\frac{S_t - M_a}{M_a - S_p}} + 1 & \text{if } 1 < d \end{cases}$$

$\omega(d) = -16.38011843$ Mars' corresponding Influenced Rotation by the Deimos on the Y axis of the graph

$$t_r := \text{if} \left(a_m < M_a, \text{if} \left(\omega_{Mi} > \omega_F, \frac{i_m \cdot \omega_F}{90}, \frac{i_m \cdot \omega_{Mi}}{90} \right), \text{if} \left(\omega(d) \cdot \omega_{Mi} > \omega_F, \frac{i_m \cdot \omega_F}{90}, \frac{i_m \cdot \omega(d) \cdot \omega_{Mi}}{90} \right) \right)$$

$t_r = -2.7924616 \times 10^{-9}$ Mars' Maximum and Free Rotational Speed Reduction by Axis Tilt

$$\omega i_3 := \text{if} \left[a_m < M_a, \omega(d) \cdot (\omega_{Mi} - t_r), \text{if} \left(q < S_t, \omega(d) \cdot \omega_{Mi} - t_r, 0 \right) \right]$$

$\omega i_3 = 0$

Mars' end result Influenced Rotation by the Deimos (p.d.)

(Negative number means the reduction amount from Mars' Free Rotation)

Part 4

Mars' Total Rotation

$$\omega_s := \sum_{i=1}^3 \omega_i$$

$$\omega_s = 0.96968927 \quad \text{Mars' Total Rotation (p.d.)}$$

$$T := \text{if} \left(\omega_1 \leq 0, 0, \text{if} \left(t \leq 90, \frac{1}{\omega_s}, \frac{-1}{\omega_s} \right) \right)$$

$$T = 1.031258 \quad \text{Mars' Sidereal Rotation Period (day)}$$

If (T = 0 , Mars' Synchronous Tropical Rotation)

Observation

$$T_o := 1.025954 \quad \text{Mars' Sidereal Rotation Period (day)}$$

If (T = 0 , Mars' Synchronous Tropical Rotation)

$$\%Diff := \frac{(T - T_o) \cdot 200}{T + T_o}$$

$$\%Diff = 0.5157 \quad \text{Percentage difference between the calculation and the observation}$$